



# MMS - Spectral Sensors

An *Extended* Introduction

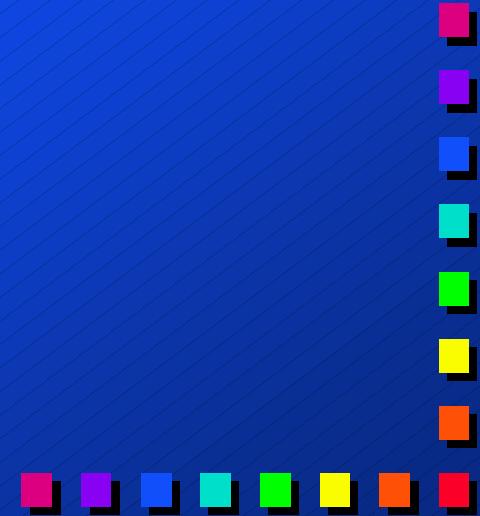


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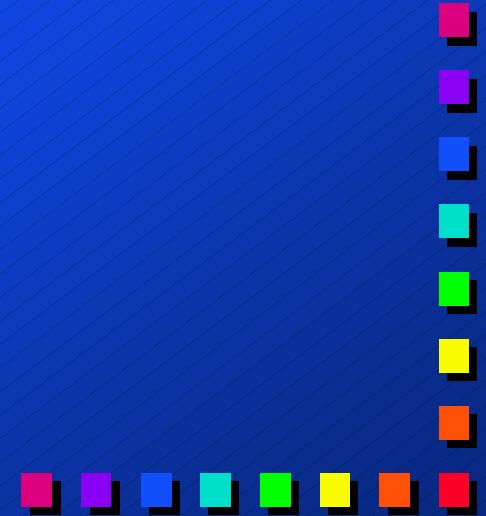
# Content

- Task
- Design and Basics
- Parts
- Features
- Overview of Current Modules
- Driving Electronics
- Forecast



# Task

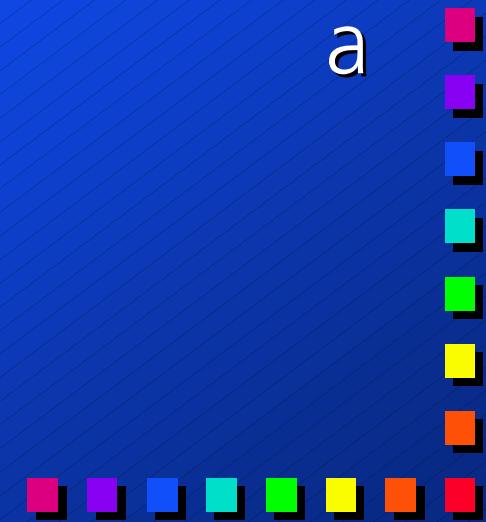
- Fast and reliable collection of complete spectra obtainable by
- Simultaneous Spectrometers



# Solution

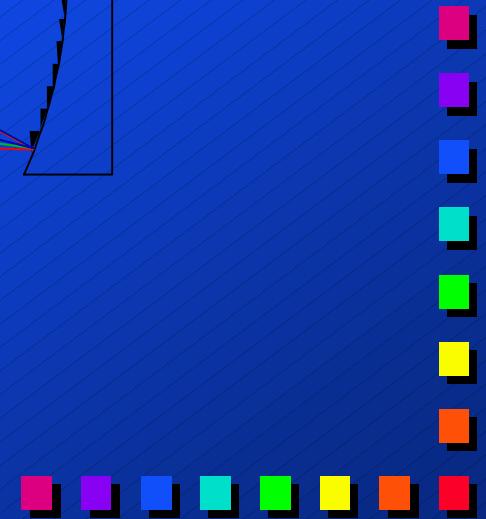
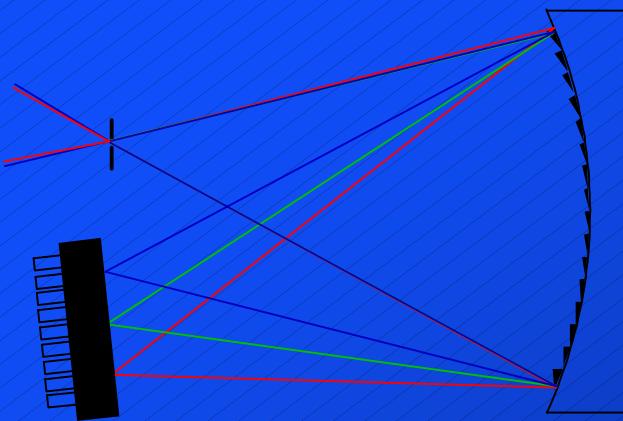
- Reliable and compact design
- ⇒
- Transformation of a classical spectrometer  
into

*Spectral Sensor*



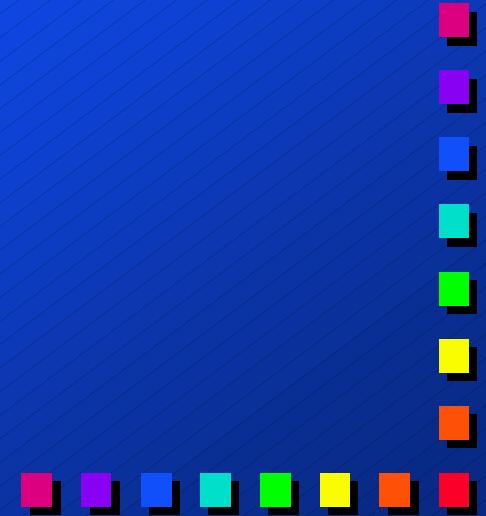
# Principal Layout

- Optical Input
- Imaging grating
- Photo diode array



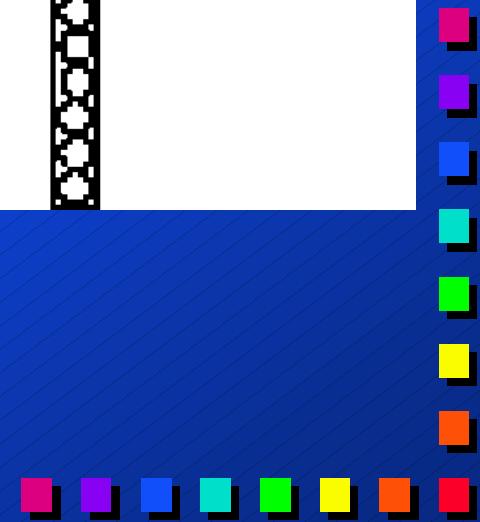
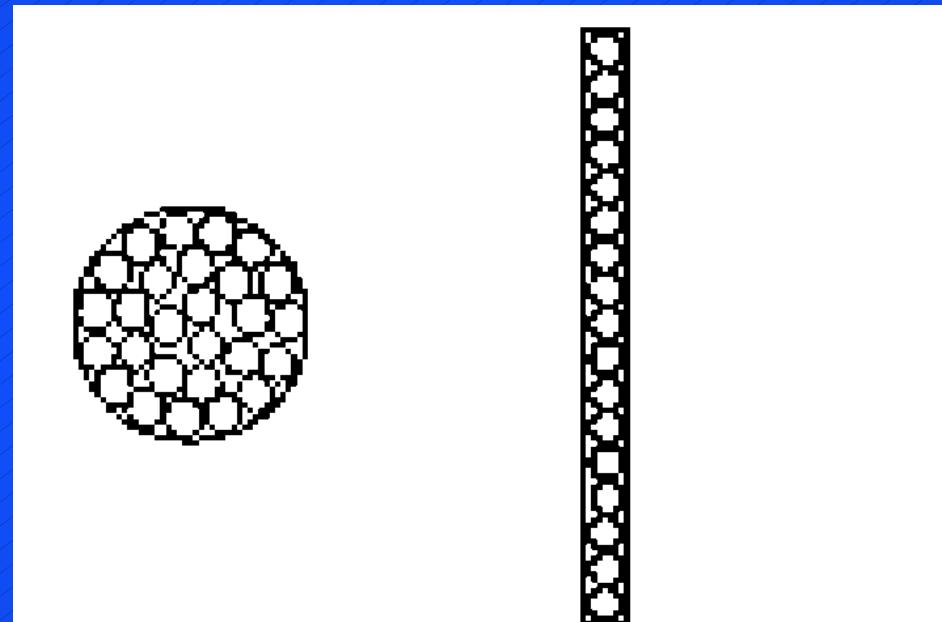
# Design Goals

- Reliable, permanently aligned
  - No movable parts
- Fast - Simultaneous read-out
- Sensitive - Optimized light throughput
- Compact



# Optical Input

- Fiber cross-section converter
- SMA connector
- NA = 0.22
- Ø = 0.5mm



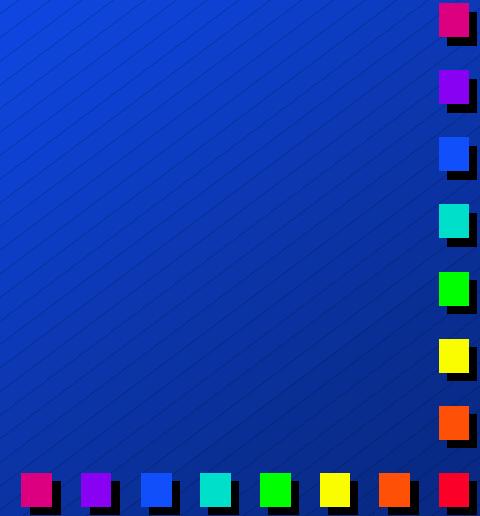
# Optical Input 2

- Width  $b_{\text{Slit}}$
- Height  $h_{\text{Slit}}$
- Effective Area  $A_{\text{eff}}$ 
  - 1:1 Imaging
- $3 \times b_{\text{Pixel}} \approx b_{\text{Slit}} = \emptyset_{\text{Fiber}}$
- $h_{\text{Pixel}} = h_{\text{Slit}} = N_{\text{Fiber}} \times \emptyset_{\text{Fiber}}$
- $A_{\text{eff}} = b_{\text{Slit}} \times h_{\text{Slit}}$



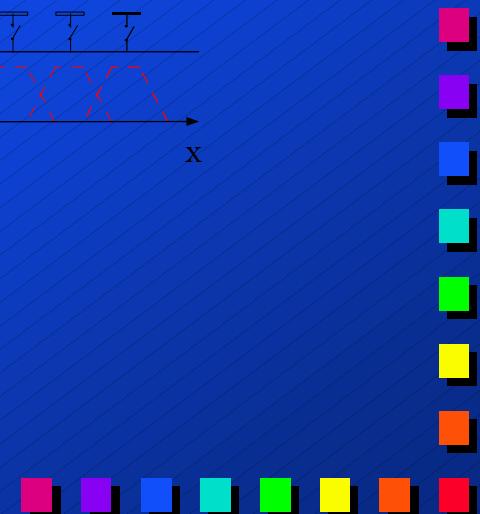
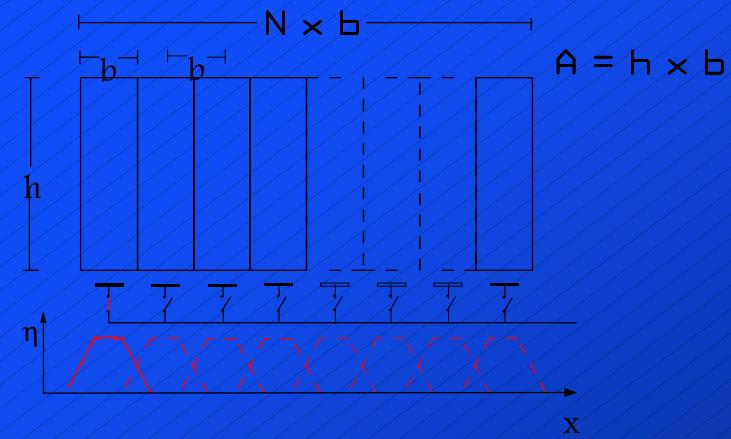
# Imaging Grating

- Holographic recording
- Flat-field design
- Imaging corrected
- Blazed = efficiency optimized
- $D/f = (2x) NA_{Fiber}$



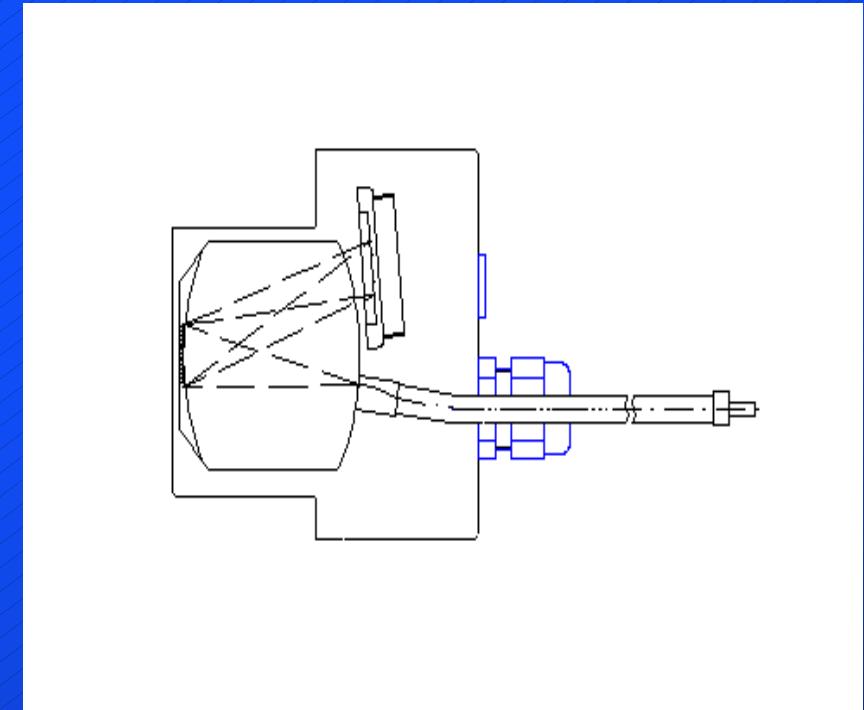
# Si Photo Diode Array

- Large pixel area
- 100% fill factor
- Pitch: 25 $\mu$ m
- Height: 2500 $\mu$ m
- Number of Pixel:  
128 ... 1024
- UV-grad\*
  - \* NIR enh. quality  
possible



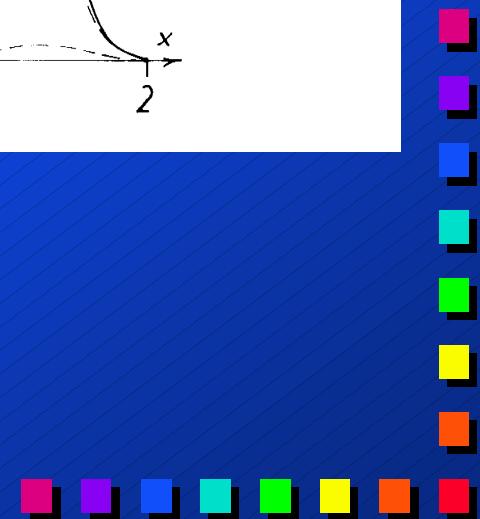
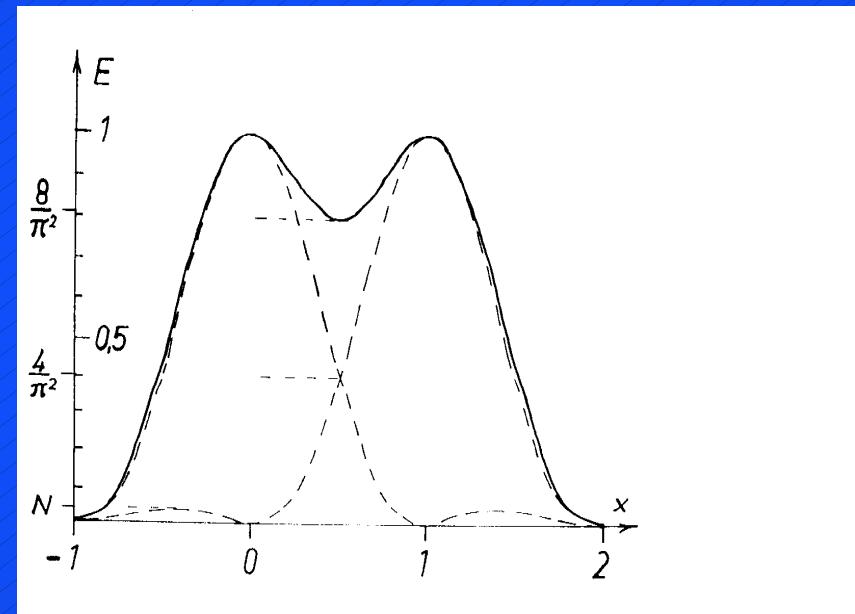
# Spectrometer

- All parts glued together
- Integrated preamplifier
- Standard interfaces



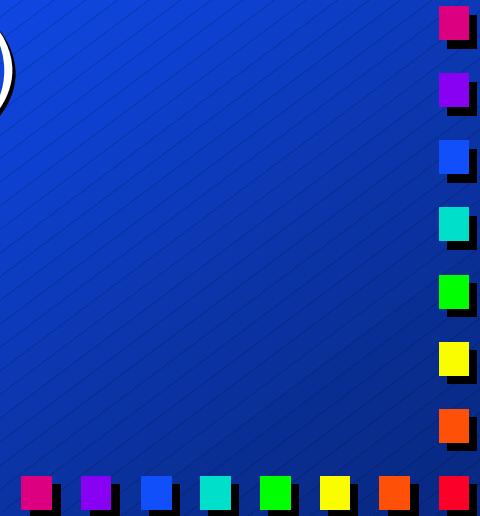
# Spectral Resolution

- Rayleigh Criterion
- FWHM
- Bandwidth
- Sub-Pixel Resolution
- Reproducibility



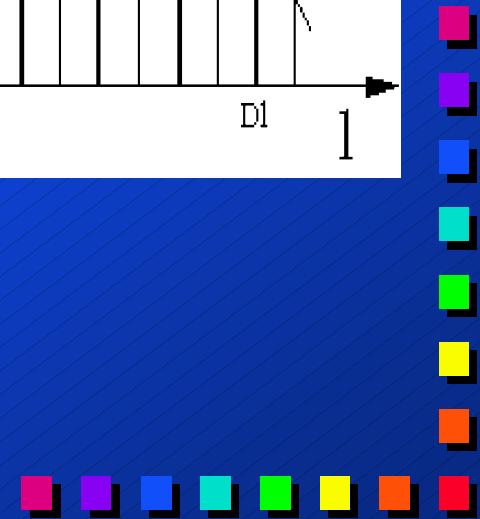
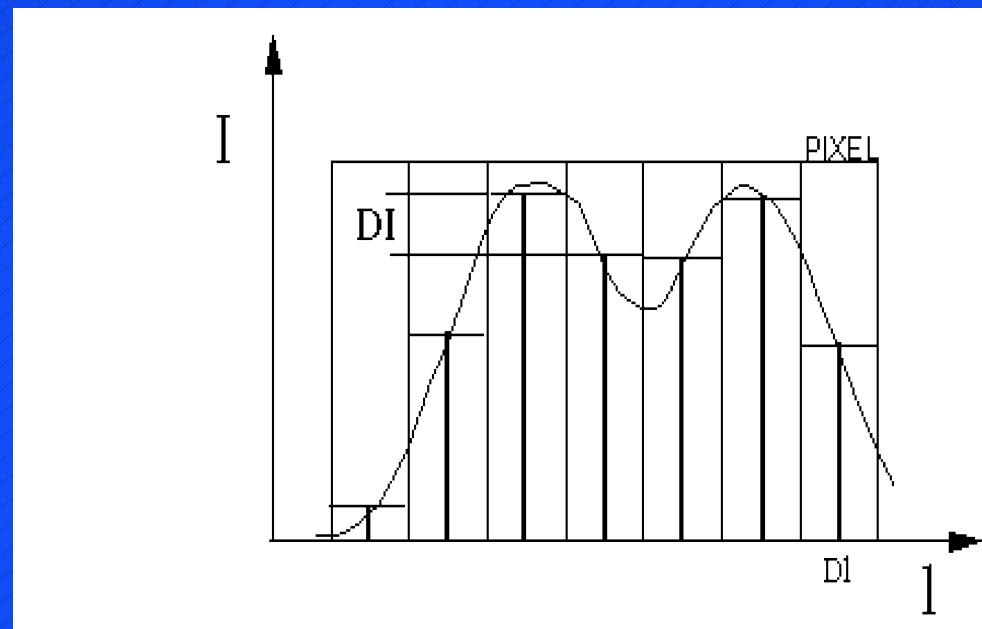
# Spectral Resolution DAS

- Rayleigh Criterion       $\Delta\lambda_{\text{Rayl.}} \approx 3 \times \Delta\lambda_{\text{Pixel}}$
- FWHM ≈  
Bandwidth       $\Delta\lambda_{\text{FWHM}} \approx 0.9 \times \Delta\lambda_{\text{Rayleigh}}$
- Sub-Pixel Res.
- Reproducibility /  
Read-Out  
Accuracy       $\Delta\lambda_{\text{SPR}} \approx 0.1 \times \Delta\lambda_{\text{Pixel}}$
- $\Delta\lambda_{\text{abs}}, (\Delta I)$



# Rayleigh Criterion DAS

- $\Delta\lambda_{\text{best case}} = 2 \times \Delta\lambda_{\text{Pixel}}$
- $\Delta\lambda_{\text{worst case}} = 3 \times \Delta\lambda_{\text{Pixel}}$



# Sub-Pixel Resolution

- Parabolic Fit:  $I(\lambda)$

$$= a \lambda^2 + b \lambda + c$$

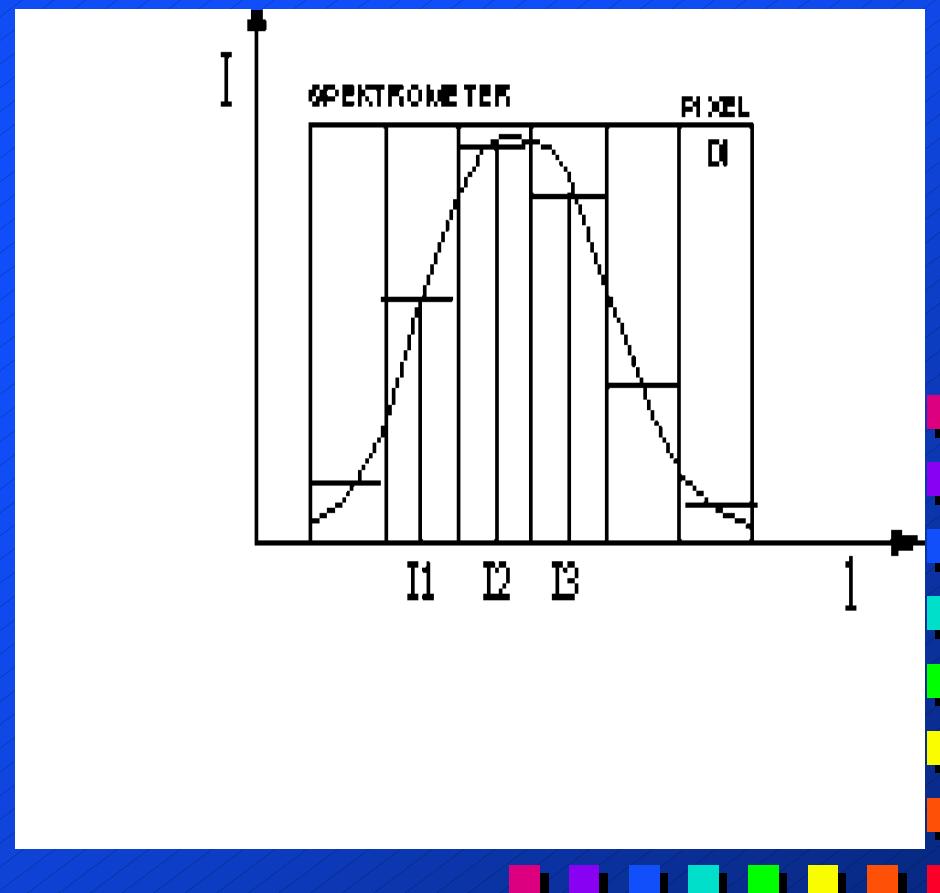
- $a = (I_3 + I_1 - 2I_2)/2\Delta\lambda^2$

- $b = (I_3 - I_1)/2\Delta\lambda - 2a\lambda_2$

- $c = I_2 - a\lambda_2^2 - b\lambda_2$

- $\lambda_{\max} = -b/2a$

- Fit Procedures for Calibration



# Intensity Resolution

- Dynamic Range
- Stability
  - = Stand. Deviation
  - = Noise
  - = Sensitivity
- Dark Current
- Linearity
- Signal / Noise
- $\Delta I = N.E.P.$
- $I_{Dark} = f(T, t_{int})$



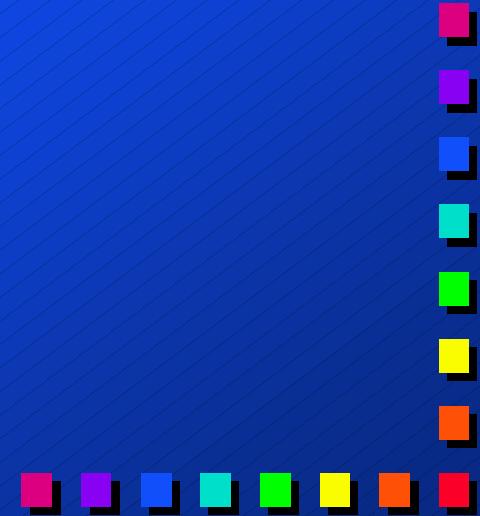
# Intensity Accuracy

- Dynamic
  - Signal / Noise
  - ADC
  - Sensitivity
  - Linearity
- False Light
  - Straylight
  - Other Orders
- S / N
- $\text{ADC} \geq \text{S/N}$
- $\Delta I = 1 \text{ Count}$
- $\Delta L < 1\%$
- Grating is Source
- Blocking Filter
- Structurization



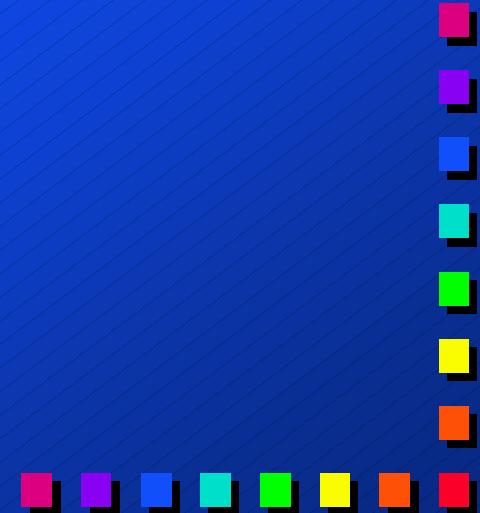
# Features

- Robust
- Permanently Calibrated
- High Dynamic Range
- High Light Sensitivity
- High Reproducibility
- Low Temperature Drift



# History

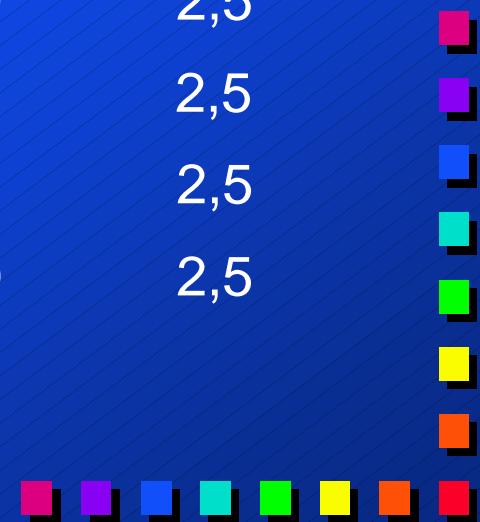
- End of 70<sup>th</sup>
  - Flat-Field-Grating + Dioden Arrays
- 1985 - MCS-Module
- 1993 - MMS 1
- 1995 - MMS UV-VIS
- 1996 - MMS UV
- 1997 - MMS NIR





# Overview Modules

Name	#Pixels	$\lambda_{\text{Start}}/\text{nm}$	$\lambda_{\text{end}}/\text{nm}$	Resol./nm	
MMS 1	256	305	1100	10	
MMS UV-VIS	256	190	740	7	
MMS UV	256	200	400	3	
MMS NIR	128	900	1680	18	
MCS UV-NIR	1024	200	1020	2,5	
MCS UV-VIS	512	200	620	2,5	
MCS VIS	512	360	780	2,5	
MCS VIS-NIR	512	680	1100	2,5	





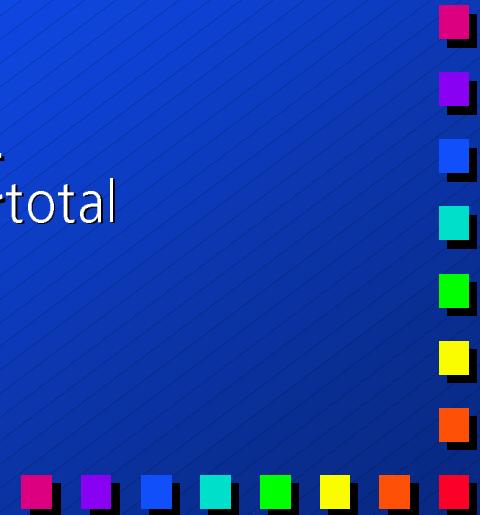
# Comparison MMS - MCS

Figure	MMS	MCS
Spectral Range	UV - NIR	UV - IR
Resolution	Low	High
Accuracy	High	Very High
Sensitivity	High	Low
False Light	Moderate	Low
Size/Weight	Small	Medium



# Driving

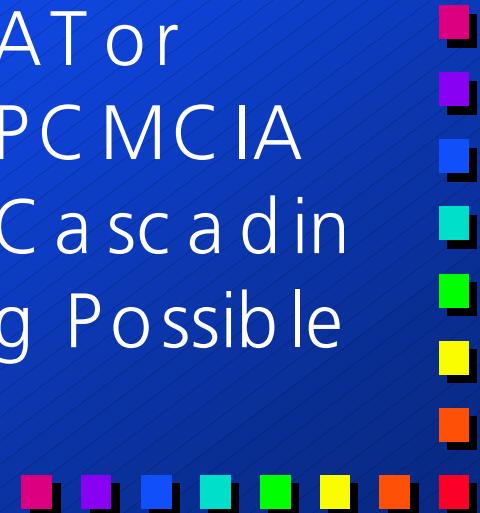
- Clock Rate  $f_{Clock}$
- Integration Time = Exposure Time
- Measurement time
- Measurement Rate (Speed)
- $t_{int,min} \geq N_{Pixel}/f_{Clock}$
- $t_{int,max}$  limited by  $I_{Dark}$
- $t_{total} = N_{cycle} \times t_{int}$
- $f_{Rate} = 1/t_{total}$





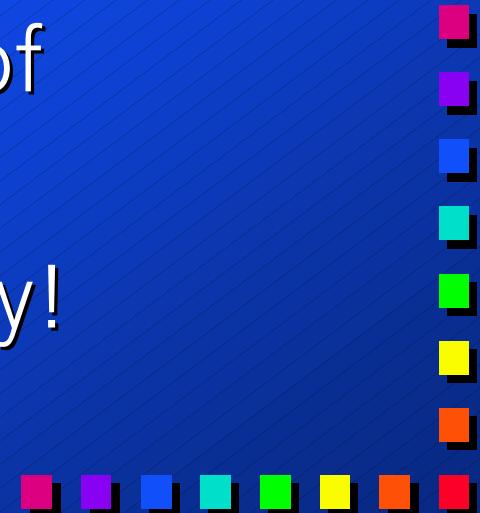
# Overview Driving Electronics

Figure	CZ 12-Bit	tec5	CZ MCS
ADC	12 Bit	15 Bit	16 Bit
Noise	1 Count	2 Counts	2 Counts
$t_{int,min}$	10 (5) ms	4ms	3ms
$t_{int,max}$	0.65s	6s	5s
Interface	XT or PCMCIA	AT PC/104	AT or PCMCIA
Multiple Modules	No	4	Cascading Possible



# Intensity Accuracy 2

- Signal should be high - Throughput!
- Besides hardware design "intelligent" software design should be applied
- Long-term stability is difficult to check
- Improvement by accumulation of measurements
- Don't forget light source stability!



# Forecast

- MMS NIR 2.4 (extented + cooled)
- MMS MIR for Mid-IR-Range
- Low-cost modules
- MCS with higher resolution

